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# The Crystal Structure of ZrAl<sub>2</sub>

BY C. G. WILSON

# Royal Military College of Science, Shrivenham, Berks., England

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The crystal structure of  $ZrAl_2$  has been determined from powder samples and imperfect single crystals. The unit cell is hexagonal with

a = 5.282, c = 8.748 Å; c/a = 1.656.

There are four  $ZrAl_2$  units per cell. The calculated density is 4.56 g.cm.<sup>-3</sup>. The space group is  $P6_3/mmc$ . The structure is a Laves phase of the  $C_{14}$ , MgZn<sub>2</sub> type. The contraction of the Al substructure and the expansion of the Zr substructure support the general conclusions about the constitution of Laves phases.

McPherson & Hanson (1954) have published the equilibrium diagram for zirconium-aluminium and reported that the intermetallic compound  $ZrAl_2$  has an orthorhombic unit cell with

$$a = 10.42, b = 7.22, c = 4.98 \text{ Å}$$
.

In order to investigate this structure more fully a sample of the alloy was prepared by arc-melting in argon and given to this laboratory by Miss J. R. Murray, A.E.R.E., Harwell. Micrographs taken from a polished and etched section showed that the sample was approximately single phase, the maximum grain size being about 0.3 mm. The impurity phase, which was estimated to be present in the approximate ratio 1:17. was later identified as the intermetallic compound Zr<sub>2</sub>Al<sub>3</sub>. Both Debye–Scherrer powder photographs and single-crystal (rotation and oscillation) photographs were used.

For the Debye-Scherrer films a precision camera of 19 cm. diameter (Adam, 1954) was kindly lent by Dr J. Adam, A.E.R.E., Harwell. The powder specimen was prepared by crushing and grinding with an agate mortar so as to pass through a  $300 \times$  sieve and inserted in a 0.3 mm. radius Lindemann-glass tube. The ZrAl<sub>2</sub> alloy is very brittle and the sharpness of the diffraction lines indicated that annealing of the powder was not necessary. All the films were taken with filtered Cu K $\alpha$ radiation. The diffraction lines were measured with an illuminated fibre device similar to that of Lees (1955) with an accuracy of  $\pm 0.005$  cm. This instrument gave much better reproducibility of readings on the weak and very weak lines than was possible with a precision travelling microscope.

From the calculated values of  $\sin^2 \theta$  it was found that all the lines could be indexed on a hexagonal unit cell with

$$a = 5.2824 \pm 0.0005$$
,  $c = 8.7482 \pm 0.0005$  Å,  
 $c/a = 1.656$ .

An extrapolation method using the Nelson-Riley

function was used to obtain accurate cell dimensions. Confirmation of this unit cell was obtained by using very small and somewhat imperfect single crystals produced by crushing a piece of the specimen. Rotation and oscillation photographs were taken with both

Table 1. X-ray data

|                     | $\sin^2	heta	imes 10^4$ | $\sin^2	heta	imes 10^4$ | Intensity        | Intensity  |
|---------------------|-------------------------|-------------------------|------------------|------------|
| hkl                 | (obs.)                  | (calc.)                 | (visual)         | (calc.)    |
| 1010                | 289                     | 287                     | w                | 10         |
| 0002                | 315                     | 315                     | vw               | 6          |
| 1011                | 367                     | 366                     | $\boldsymbol{w}$ | 8          |
| 1012                | 601                     | 601                     | m-               | 10         |
| $11\bar{2}0$        | 859                     | 858                     | 8                | 40         |
| 1013                | 992                     | 990                     | vs               | <b>62</b>  |
| $20\overline{2}0$   | 1143                    | 1143                    | m-               | 9          |
| $11\overline{2}2$   | 1170                    | 1170                    | vs               | 61         |
| $20\overline{2}1$   | 1220                    | 1221                    | m                | 32         |
| 0004                | 1250                    | 1251                    | vvw              | 3          |
| $10\overline{1}4$   | 1535                    | 1535                    | vvw              | 2          |
| $20\overline{2}3$   | 1843                    | 1844                    | w                | 6          |
| $21\overline{3}0$   | 1995                    | 1996                    | vw               | 2          |
| $21\overline{3}1$   | 2071                    | 2075                    | vw               | 3          |
| $10\overline{1}5$   | 2232                    | 2235                    | m-               | 13         |
| $21\overline{3}2$   | 2306                    | 2307                    | vw               | 4          |
| $30\bar{3}0$        | 2562                    | 2565                    | m-               | 12         |
| $21\overline{3}3$   | 2694                    | 2696                    | 8                | <b>4</b> 0 |
| 0006                | 2806                    | 2806                    | vvw              | 4          |
| 30 <u>3</u> 2       | 2872                    | 2871                    | m                | 21         |
| 20 <u>2</u> 5 )     | 3083                    | ∫ 3082                  | m                | 32         |
| 10 <u>1</u> 6 ∫     | 0000                    | 3085                    |                  | 02         |
| 2134                | 3237                    | 3235                    | vvw              | 2          |
| 2240                | 3413                    | 3410                    | m                | 22         |
| $11\overline{2}6$   | 3654                    | 3652                    | vvw              | <b>2</b>   |
| $22\overline{4}2$   | 3724                    | 3721                    | vvw              | 3          |
| 21 <del>3</del> 5 ) | 2025                    | f 3932                  | <b>1</b> 03      | 99         |
| 2026 ∫              | 0900                    | <b>)</b> 3935           | m                | 22         |
| $31\overline{4}2$   | 4003                    | 4005                    | vw               | <b>2</b>   |
| $10\overline{1}7$   | 4096                    | 4092                    | vvw              | 3          |
| $31\overline{4}3$   | 4392                    | <b>4392</b>             | m                | 24         |
| $40\overline{4}0$   | 4548                    | 4544                    | vvw              | <b>2</b>   |
| $40\bar{4}1$        | 4622                    | 4622                    | vw               | 7          |
| $21\overline{3}6$   | 4783                    | 4784                    | vvw              | 5          |
| 40 <del>4</del> 3)  | 5940                    | 5242                    |                  | 0          |
| 1018                | 0249                    | 5254                    | vw               | ฮ          |
| $30\overline{3}6$   | 5351                    | 5351                    | vvw              | 2          |
| $31\overline{4}5$   | 5631                    | 5631                    | w                | 12         |

|   | $\sin^2	heta	imes 10^4$ | $\sin^2	heta	imes 10^4$                  | Intensity | Intensity |
|---|-------------------------|--|-----------|-----------|
| hkl   | (obs.)                  | (calc.)                                  | (visual)  | (calc.)   |
| $\left. \begin{array}{c} 40\overline{4}4\\ 21\overline{3}7 \end{array} \right\}$          | 5789                    | $\begin{cases} 5784 \\ 5791 \end{cases}$ | vvw       | 5         |
| $11\overline{2}8$   | 5820                    | 5821                                     | m-        | 13        |
| $41\overline{5}0$   | 5962                    | 5961                                     | m-        | 13        |
| $\left. \begin{smallmatrix} 32\overline{5}3\\ 20\overline{2}8 \end{smallmatrix} \right\}$ | 6092                    | { 6091<br>6092                           | m-        | 25        |
| $22\overline{4}6$   | 6202                    | <b>6200</b>                              | m-        | 13        |
| $41\overline{5}2$   | 6271                    | 6270                                     | m         | <b>26</b> |
| $40\overline{4}5$   | 6483                    | 6481                                     | m         | 16        |
| $10\overline{1}9$   | 6569                    | 6570                                     | vvw       | 3         |
| $21\overline{3}8$   | 6953                    | 6953                                     | m-        | 15        |
| $32\overline{5}5$   | 7332                    | 7330                                     | m-        | 15        |
| $30\overline{3}8$   | 7518                    | 7518                                     | m-        | 17        |
| $33\overline{6}0$   | 7660                    | 7659                                     | w         | 9         |
| $50\overline{5}3$   | 7789                    | 7789                                     | m-        | 15        |
| $42\overline{6}0$   | 7938                    | 7942                                     | vvw       | 6         |
| $33\overline{6}2$   | 7970                    | 7969                                     | m-        | 18        |
| $42\overline{6}1$   | 8016                    | 8019                                     | m-        | 17        |
| $32\overline{5}6$   | 8183                    | 8181                                     | w         | 7         |
| $21\overline{3}9$   | 8268                    | 8268                                     | w         | 8         |
| $22\overline{4}8$   | 8368                    | 8367                                     | vvw       | 5         |
| $\left. \begin{array}{c} 42\overline{6}3\\ 31\overline{4}8 \end{array} \right\}$          | 8647                    | $\begin{cases} 8638 \\ 8650 \end{cases}$ | m         | 34        |
| $41\overline{5}6$   | 8747                    | 8747                                     | m-        | 11        |
| $2,0,\overline{4},10$   | 8889                    | 8889                                     | w         | 13        |
| $50\overline{5}5$   | 9028                    | 9027                                     | m-        | 14        |
| $32\overline{5}7$   | 9188                    | 9187                                     | w         | 13        |
| $51\overline{6}3$   | 9487                    | 9487                                     | vs        | 74        |
| 1,0,1,11  | 9665                    | 9665                                     | m         | 27        |
|   |                         | •  | ,         | ,         |

Table 1 (cont.)

vvw = very very weak, vw = very weak, w = weak, m - = medium weak, m = medium strong, s = strong, vs = very strong. All  $sin^2 \theta$  values below dotted line refer to resolved

 $\alpha_1$  components.

a and c as axes of rotation. The observed and calculated values of  $\sin^2 \theta$  are given in Table 1.

Using the atomic volumes of the zirconium and aluminium atoms derived from the elementary crystalline forms, the proposed unit cell contains four units of ZrAl<sub>2</sub>. The calculated density was 4.56 g.cm.<sup>-3</sup>; the density determined experimentally, using a bulk sample containing the impurity phase, was 4.42 g.cm.<sup>-3</sup>.

The visually estimated and calculated line intensities are given in Table 1. The calculated intensities were corrected for absorption but the temperature factor was ignored. The absorption factor was obtained from a table given in Klug & Alexander (1954) for cylindrical powder specimens. The value of  $\mu r = 3 \cdot 1$ , used for this purpose, was based on a calculated composite absorption coefficient for the compound and an experimentally determined powder density.

The single-crystal photographs indicated the following conditions for reflection: (000l), l=2n;  $(hh2\bar{h}l)$ , l=2n; (hkil), for h-k=3n, l=2n. Although several space groups are possible a good intensity agreement was found using  $P6_3/mmc$  with the atoms in the following positions:

4 Zr in (f) positions: 
$$\frac{1}{3}, \frac{2}{3}, z; z = \frac{1}{16}$$
  
6 Al in (h) positions:  $x, 2x, \frac{1}{4}; x = -\frac{1}{6}$   
2 Al in (b) positions:  $0, 0, 0.$ 

No attempt was made to refine the Zr and the Al parameters. The proposed  $ZrAl_2$  structure is thus a Laves phase of the  $C_{14}$ , MgZn<sub>2</sub> type. The interatomic distances are given in Table 2.

| Zr-Zr: | 2 at 3.28  |   | and 2 at 3·24 Å              |       |
|--------|------------|---|------------------------------|-------|
| Al-Al: | 6 at 2.67  |   | and 2 at $2.64$ Å            |       |
| Zr-Al: | 1 at 3.13, | 1 | at $3 \cdot 11$ Å and 1 at 3 | ∙08 Å |

#### Discussion

In the ideal Laves phase of the  $C_{14}$  type which is formed when  $c/a = (8/3)^{\frac{1}{2}} = 1.633$ , the *B* atoms occupy corners of regular tetrahedra which are joined alternately point to point and base to base. The larger *A* atoms occupy the 'holes' formed by the tetrahedra to form another substructure which is characteristic of the phase. Any departure from  $c/a = (8/3)^{\frac{1}{2}}$ , leads to a change in the height of the *B* tetrahedra and consequent small differences in the interatomic distances in each substructure as observed in Table 2.

The important parameter governing the formation of Laves phases, when the c/a ratio is approximately correct, is the ratio of the atomic diameters of the two components  $d_A/d_B$ . In the ideal C<sub>14</sub> structure, in which A atoms touch each other and B atoms touch each other,  $d_A/d_B = (1\cdot 5)^{\frac{1}{2}} = 1\cdot 225$ . Berry & Raynor (1953) showed that if  $W = d_A/d_B$  is less than 1.225, there should be a contraction of the B substructure and an expansion of the A substructure. Assuming the atomic diameters  $d_A$  and  $d_B$  to be the interatomic distances in pure metallic zirconium and aluminium, 3.17 and 2.862 Å respectively, we find for ZrAl<sub>2</sub> that W = 1.108. Although this is a rather low value there are at least two other compounds, viz. WFe2 and  $MnBe_2$ , for which W is smaller. On the other hand, it is interesting to note that there is no Laves phase of the  $C_{14}$  type with Al as the B atom reported in the list given by Berry & Raynor. The expected contraction of the B substructure and the expansion of the A substructure is clearly revealed by the interatomic distances given in Table 2.

Following Berry & Raynor, and using the maximum interatomic distances listed in Table 2, the parameters  $S_{AA}$ ,  $S_{BB}$  and  $S_{AB}$ , are determined for ZrAl<sub>2</sub>. These parameters represent the percentage deviation from the interatomic distances in the pure components and are defined by the expressions:

$$\begin{split} S_{AA} &= 100(d_{AA} - d_A)/d_A, \quad S_{BB} &= 100(d_{BB} - d_B)/d_B , \\ S_{AB} &= \frac{100[d_{AB} - \frac{1}{2}(d_A + d_B)]}{\frac{1}{2}(d_A + d_B)}. \end{split}$$

For ZrAl<sub>2</sub>:  $S_{AA} = 3.48\%$ ,  $S_{BB} = -6.84\%$  and  $S_{AB} = 3.79\%$ . The values of  $S_{AA}$ ,  $S_{BB}$  and  $S_{AB}$  for the known  $C_{14}$  Laves phases formed by zirconium as the A component are plotted against W in Fig. 1, to give a diagram similar to that first presented by Berry &



Fig. 1. Plot of  $S_{AA}$ ,  $S_{BB}$  and  $S_{AB}$  against  $W = d_A/d_B$  for Laves phases of the  $C_{14}$  type with zirconium as the 'A' component.

 $----S_{BB}; ----S_{AB}.$ 

Raynor. Aluminium appears to fit very well into this scheme and since chemically aluminium differs considerably from all other metals in this set, the importance of geometrical factors in the formation of the  $C_{14}$  phase seem to be over-riding.

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## References

ADAM, J. (1954). J. Sci. Instrum. 31, 131.

- BERRY, R. L. & RAYNOR, G. V. (1953). Acta Cryst. 6, 178.
- KLUG, H. P. & ALEXANDER, L. E. (1954). X-ray Diffraction Procedures, p. 155. New York: Wiley.
- LEES, C. S. (1955). J. Sci. Instrum. 32, 17.
- McPHERSON, D. J. & HANSON, M. (1954). Trans. Amer. Soc. Metals, 46, 354.

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# Transition Element—Rare Earth Compounds with the Cu<sub>5</sub>Ca Structure

By J. H. WERNICK AND S. GELLER

Bell Telephone Laboratories, Inc., Murray Hill, New Jersey, U.S.A.

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A number of new  $B_5A$  compounds, with A a rare earth or yttrium and B a transition element, Co, Ni, or Cu, having the Cu<sub>5</sub>Ca structure, have been prepared. In most of the compounds, the interatomic distances are normal. However, Ce appears to have a valence substantially greater than three. Also, in Co<sub>5</sub>Pr, Pr appears to have a valence slightly greater than three.

## Introduction

In the course of a continuing study of the magnetic and structural properties of intermetallic compounds between rare earth and transition metals, a number of new compounds with the Cu<sub>5</sub>Ca structure have been made. Some of the results of magnetic measurements have already been reported by Nesbitt *et al.* (1959). It is the purpose of this paper to report on the crystallographic investigations of these new compounds.

#### Experimental

Stoichiometric amounts of the constituents were melted by induction heating in fused  $Al_2O_3$  or quartz

crucibles in an argon atmosphere.\* The compounds appear to melt congruently as they form readily from the melt.

X-ray powder photographs were taken with Cr  $K\alpha$  radiation and a Straumanis type Norelco camera (114.6 mm. diameter). Observed and calculated interplanar spacings are shown in Table 1.

These compounds,  $B_5A$ , belong to space group  $P6/mmm(D_{6h}^1)$  with A in (a): 0, 0, 0,  $2B_{\rm I}$  in (c):  $\pm (\frac{1}{3}, \frac{2}{3}, 0)$ , and  $3B_{\rm II}$  in (g):  $\frac{1}{2}, 0, \frac{1}{2}, 0, \frac{1}{2}, \frac{1}{2}, \frac{1}{2}, \frac{1}{2}, \frac{1}{2}, \frac{1}{2}$ 

<sup>\*</sup> All of the rare earth elements, except Pr (99.9+%), and Y were of 99+% purity. Co and Ni 99.9+%; Cu 99.999%. The Nd was kindly supplied to us by Dr F. H. Spedding.